

Energy filtered PEEM

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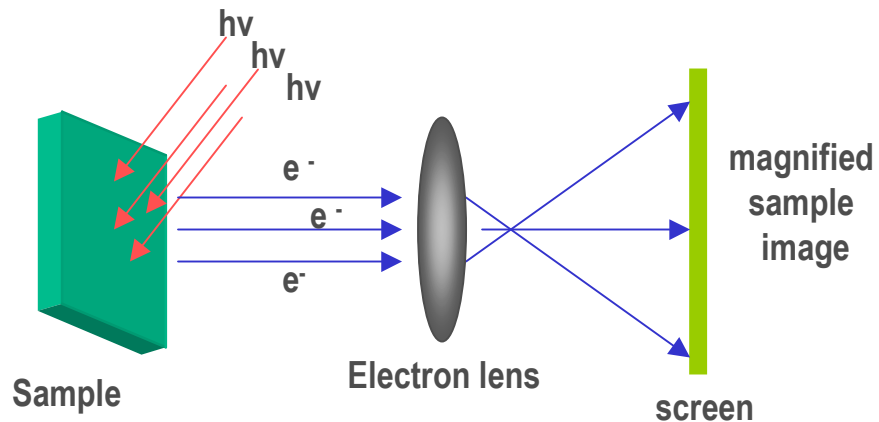
The Principle of PEEM

The Photoelectron Emission Microscope

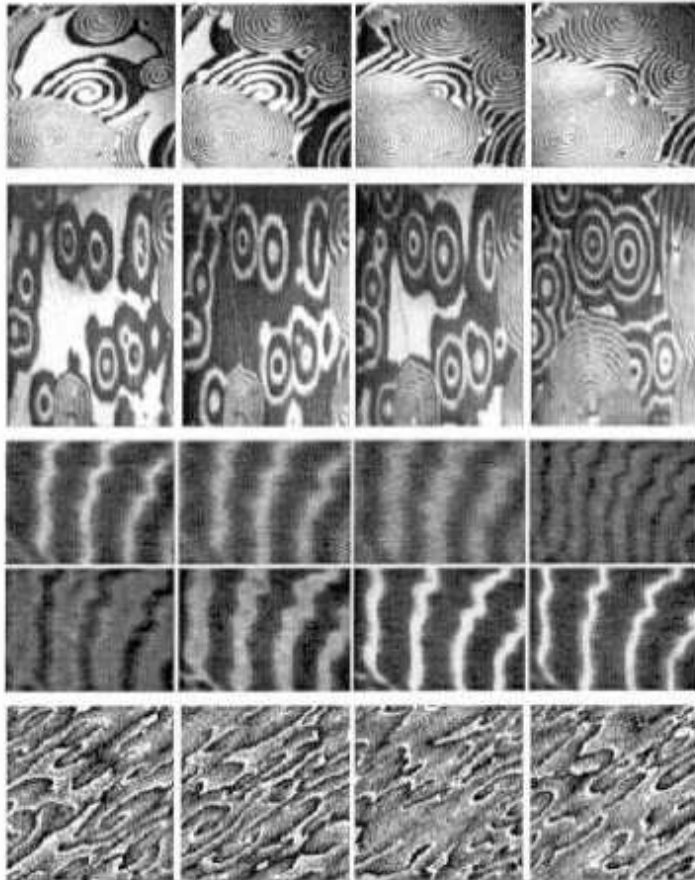
surface sensitive to the top most layer

non destructive

imaging lens system allows for massive parallel detection



A brief History



Spiral patterns

catalysis research:

The PEEM images show the formation of spiral wave patterns in catalytic CO oxidation on a clean Pt(110) surface

Chaos

Progression of time => => => => => => =>

Platinum surface imaged by photoemission electron microscopy. Dark areas are rich in CO while light areas are O₂ rich. Note the oscillatory behavior of the domain extensions. Time scale ~10s, length scale ~0.1mm (The Surface Imaging Group, Dept. of Physical Chemistry, Fritz-Haber-Institute of the Max-Planck-Society, www.fhi-berlin.mpg.de/surfimag)

A brief History

- 1933 invented by Brüche
- 1971 Metioskope KE3 from Balzers
- 1970 -1990s PEEM has been utilised by several groups

- Prominent is example Ertl: Nobel Prize in chemistry 2007

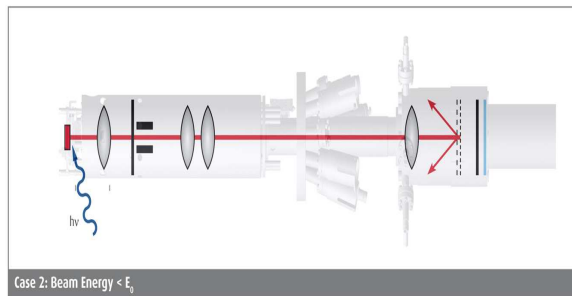
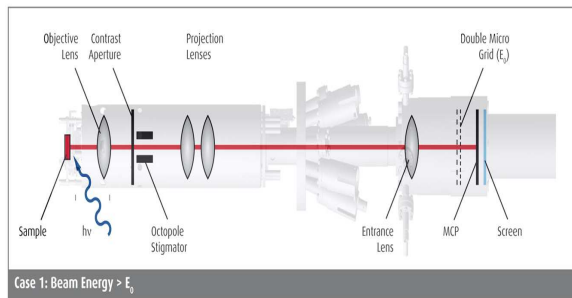
- around 1988: Tonner et al.: first Energy filtered PEEM
- followed by Ernst Bauer et al. in the early 90's

- Late 90's several Energy filtered PEEM concepts have been commercialised

Energy filtered PEEM

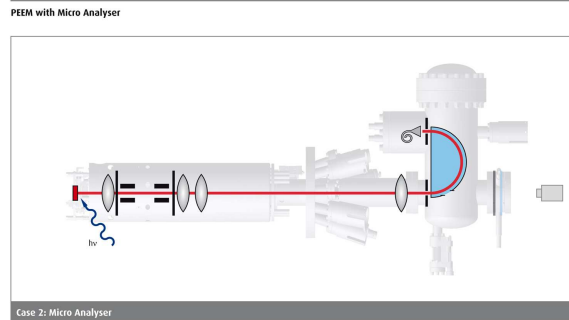
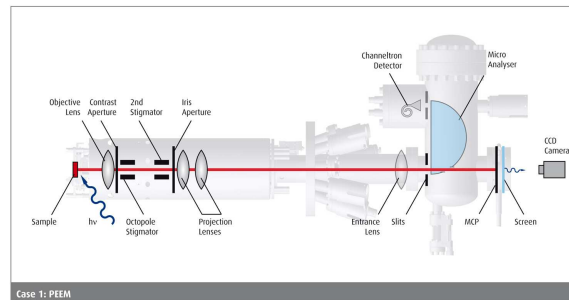
- High Pass imaging energy filter,
- Micro Spot electron spectroscopy,
- Band Pass imaging analyser

PEEM with Imaging High-Pass Energy Filter

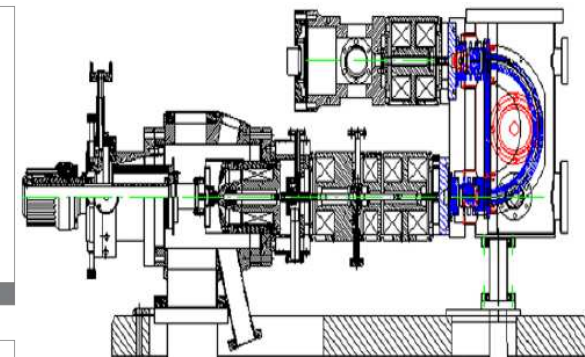


Staib
Focus

PEEM with Micro Analyser



Focus



Elmitec

Such energy filters demand for very high photon flux densities when working with higher kinetic electron energies $> 10\text{eV}$

Lab Experiments: $E_{\text{kin}} < 10\text{eV}$

- Local work function analysis / Field emission
- Laser experiments / Plasmon
- Structure analysis / Growth processes

Synchrotrons:

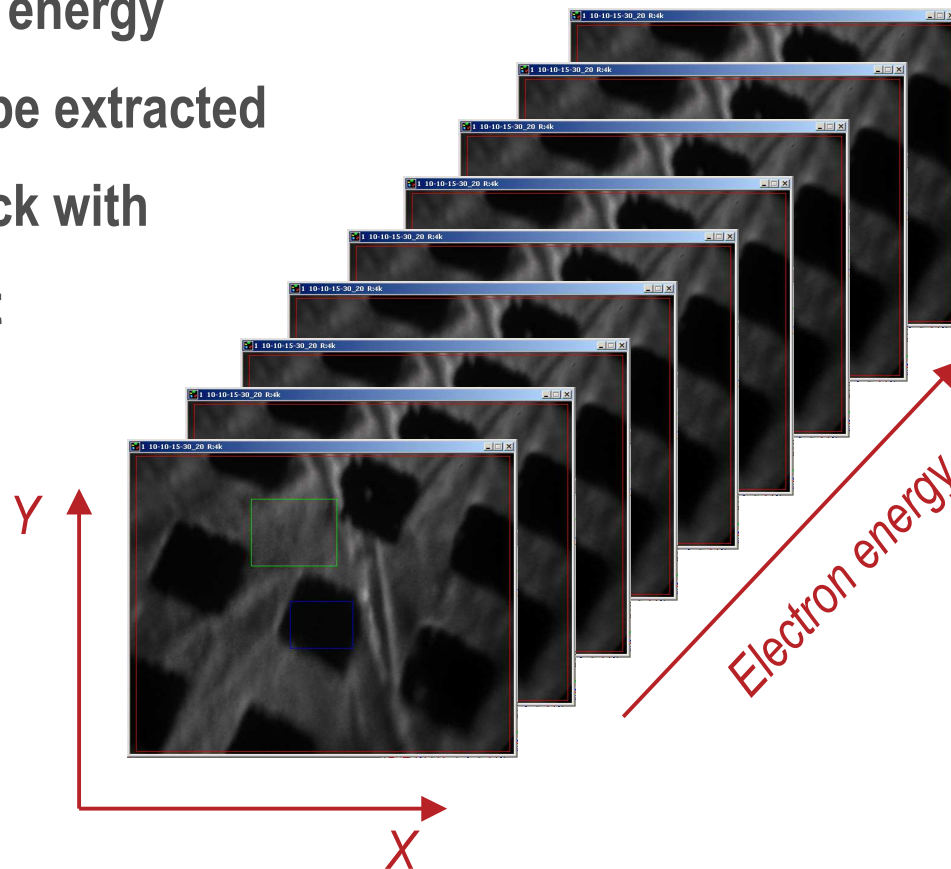
XAS experiments and XMCD using pol. Light $E_{\text{kin}} \ll 50\text{eV}$

Only limited nr. of XPS experiments $E_{\text{kin}} > 50\text{eV}$

Data Acquisition:

Multi-Dimensional Data Sets (Intensity $I(X, Y, E_{kin})$)

- The Energy filtered PEEM image is a local area map at single energy
- Local spectra can be extracted from image an stack with images recorded at different energies



Data Analysis:

Sample: Ag evaporated on Ta, $h\nu = 700$ eV

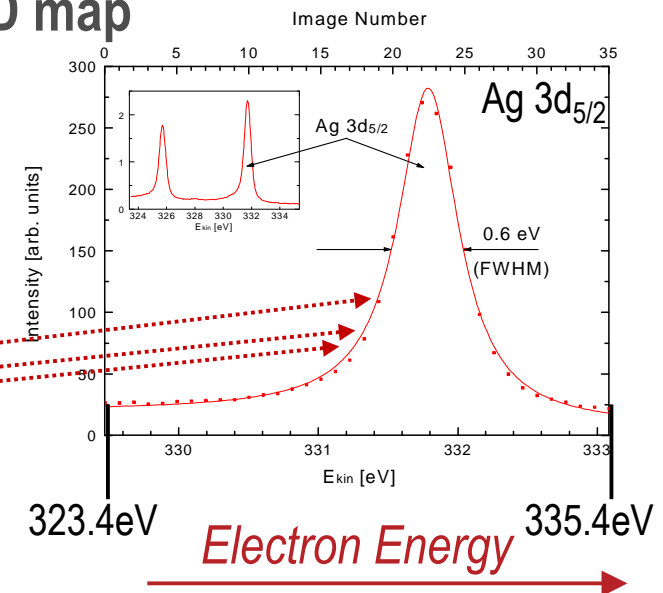
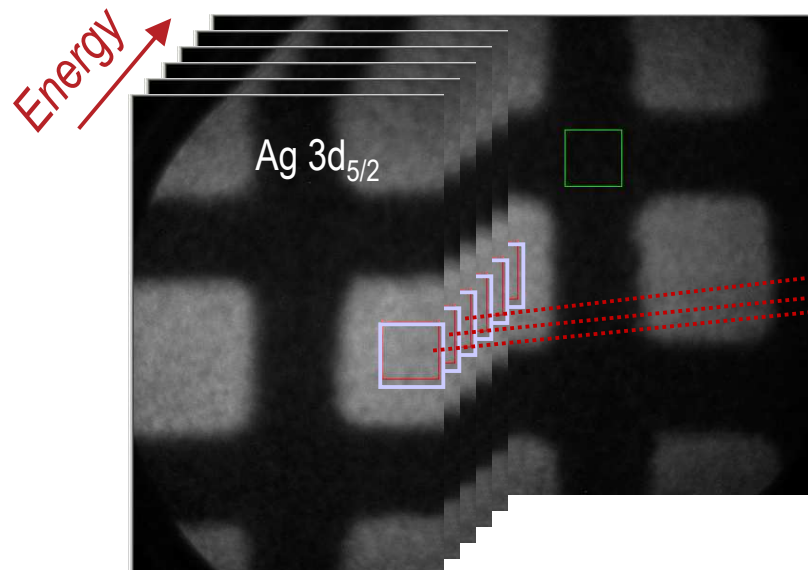
- Image stack:

36 images across Ag 3d, field of view of $35 \mu\text{m}$

- Spectrum

“Grey level” averaging within a $3 \times 3 \mu\text{m}$ area

Each spectral point is extracted from a 2D map



Aberration correctors in Photoemission Microscopy

Recent PEEM development: two main directions

A) The mirror corrector

- + ultimate PEEM / LEEM resolution in the sub 10nm range
- + increased transmission
- increased complexity of the instruments
- limits of the energy filter remain

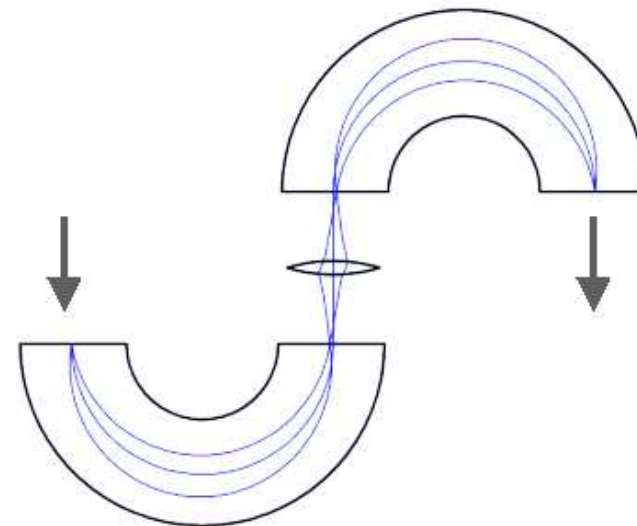
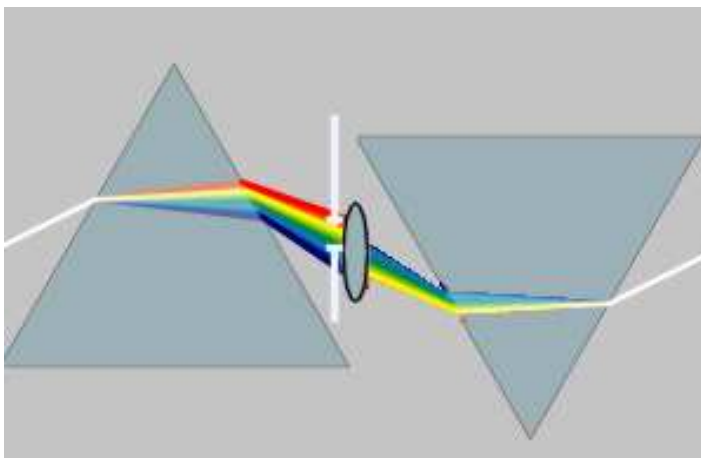
Bessy: SMART	operational
ALS PEEM III	ongoing commissioning
Commercial solutions	soon available

Aberration correctors in Photoemission Microscopy

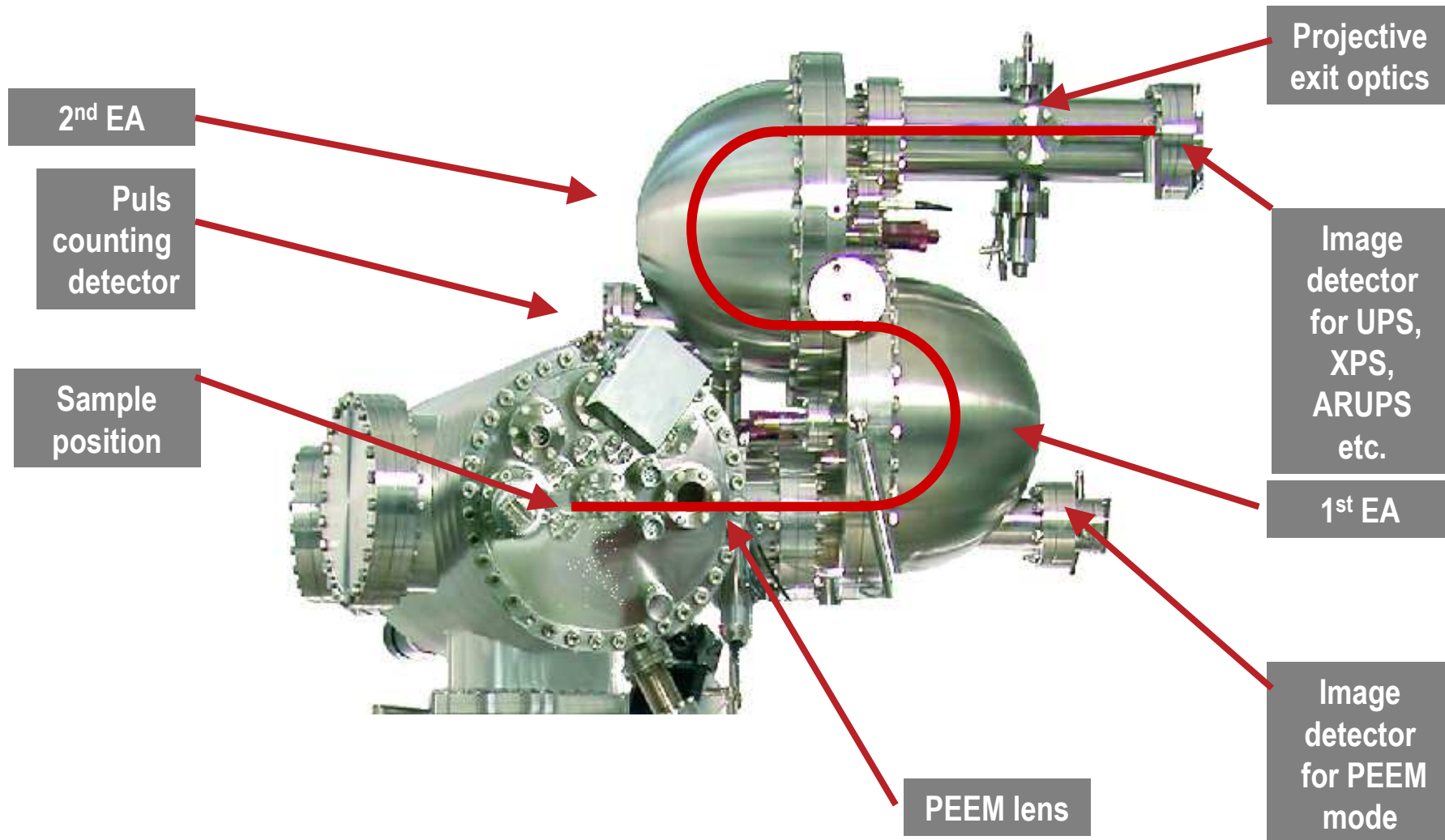
B) The aberration corrected hemispherical energy filter

Aberrations introduced by the first analyser are compensated by second analyser

- + very high transmission (larger slits)
- + excellent energy resolution (smaller pass energy)
- no increase in lateral resolution

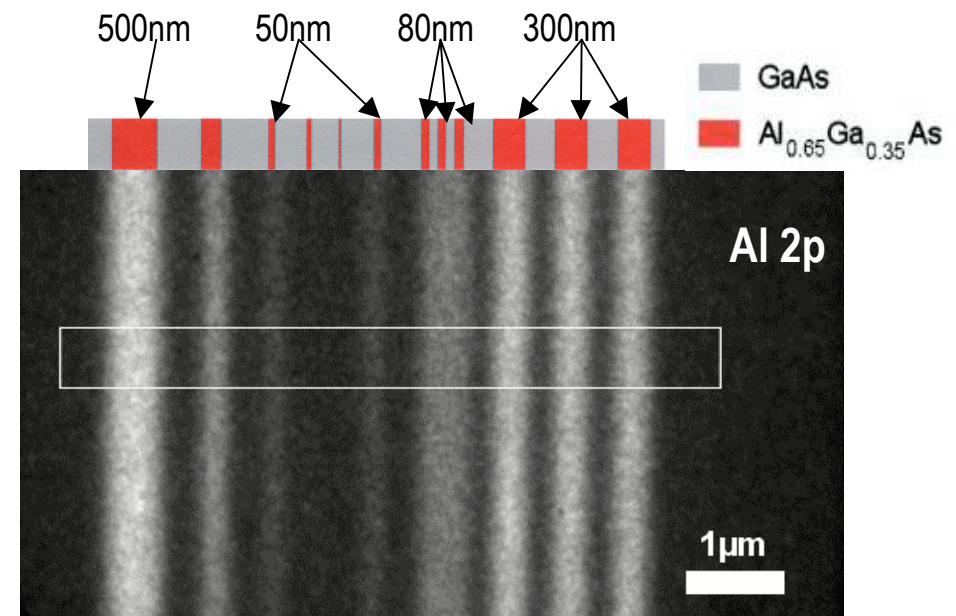


NanoESCA



Synchrotron Results (BESSY II UE 52 SGM)

- Imaging Al 2p Corelevel
- Calibrated sample from BAM
- Semiconductor heterostructure GaAs/AlGaAs (BAM L002)
- Beamline UE 52 SGM
- $h\nu = 150 \text{ eV}$, $E_{\text{kin}} = 77 \text{ eV}$
- Edge resolution
120nm @ 57% Contrast

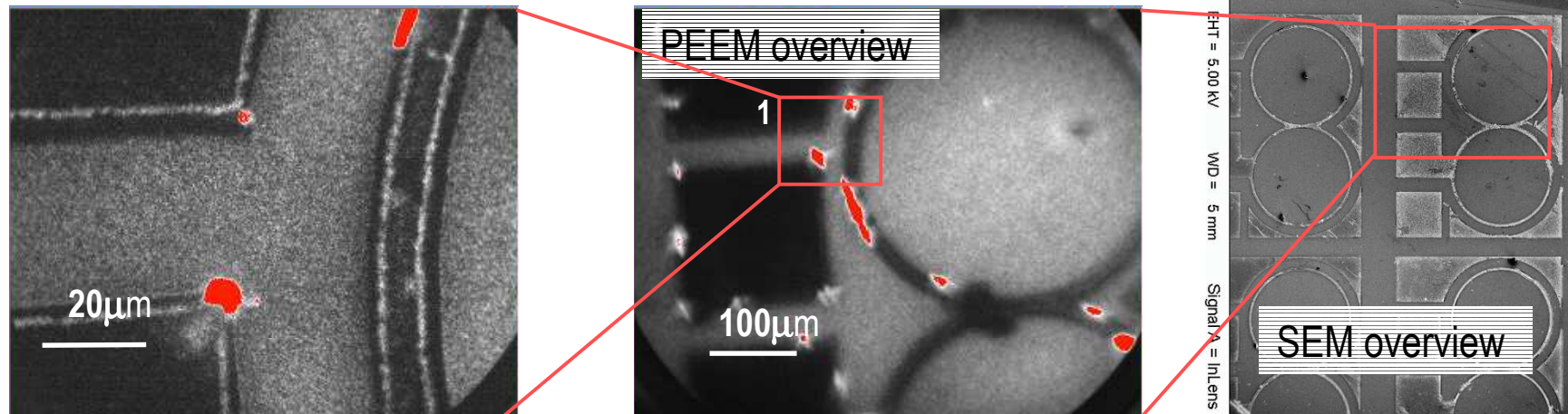


M. Senoner et al. *Journal of Surface Analysis* **12**, pp 78-82 (2005).

Synchrotron Results (ESRF ID08)

Sample: electrochemical coated with pMAN (polymethacrylonitril)

- Mercury arc lamp illumination
- PEEM mode: find good area with little contamination (red) - fast real time imaging

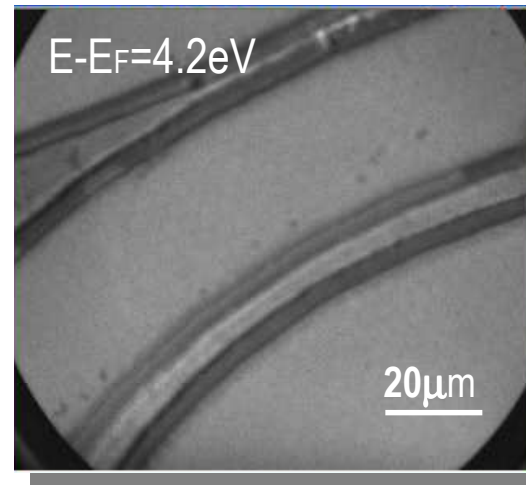
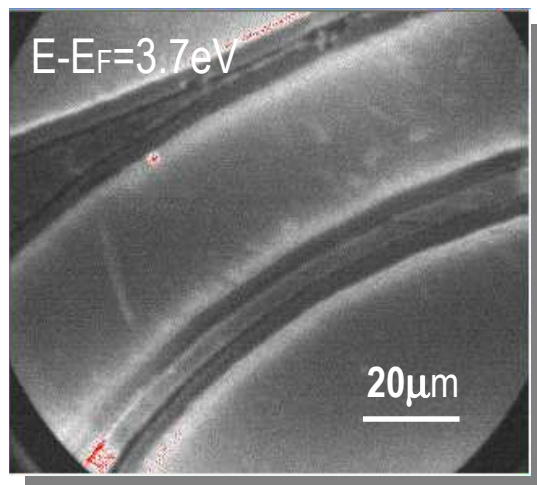


Synchrotron Results (ESRF ID08)

pMAN sample

- SE imaging with Work- function contrast
- Contrast changes selected analyser energy
- Strong indication for localised chemical compounds

Synchrotron illumination $h\nu = 700\text{eV}$

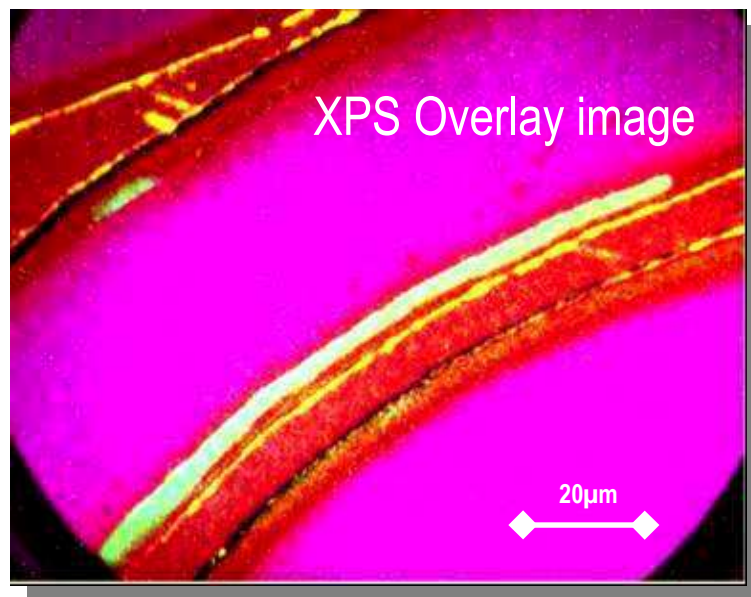


Synchrotron Results (ESRF ID08)

Understanding complex chemical map

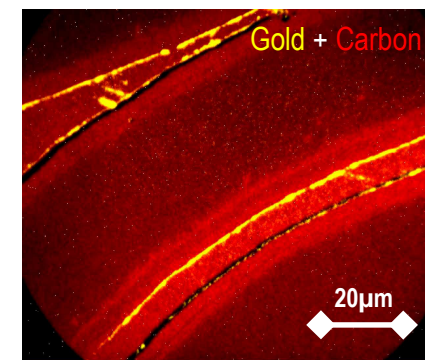
XPS chemical state mapping O(x)

$h\nu = 700\text{eV}$, ~5 min per image



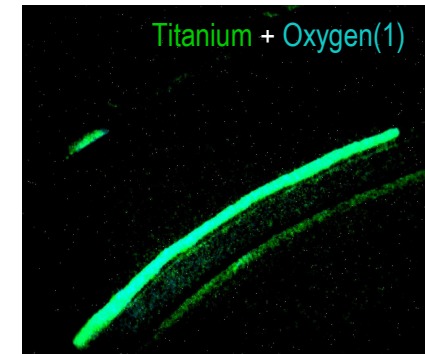
Au 4f_{7/2}, 84 eV

C 1s, 284.2 eV

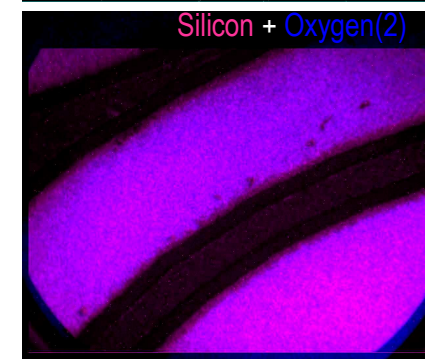


Ti 2p_{3/2}, 453.8 eV

O 1s, 543.1 eV



Si 2p_{3/2}, 99.4 eV



O. Renault et al., *Surface Science*. 601, 2007

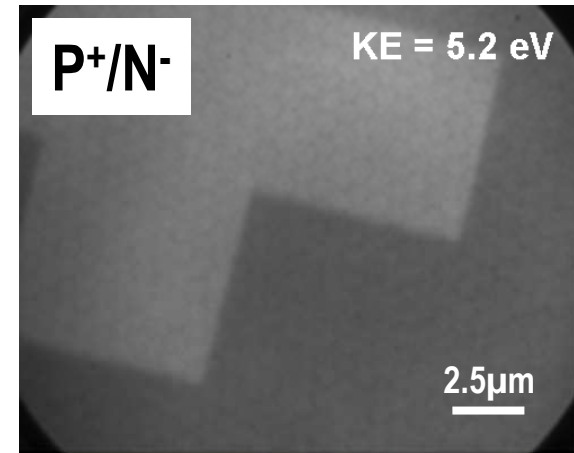
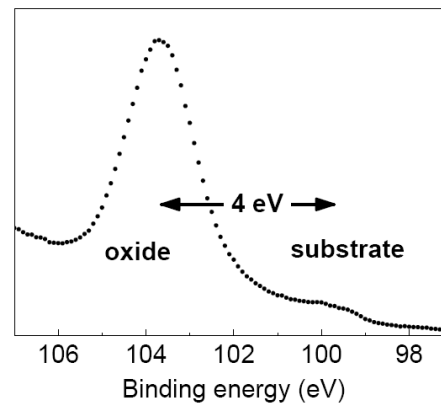
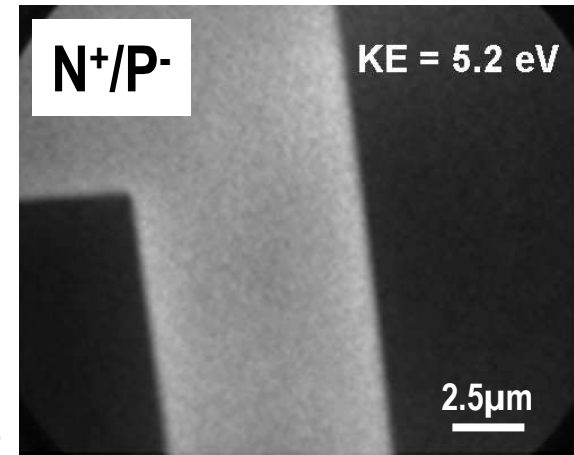
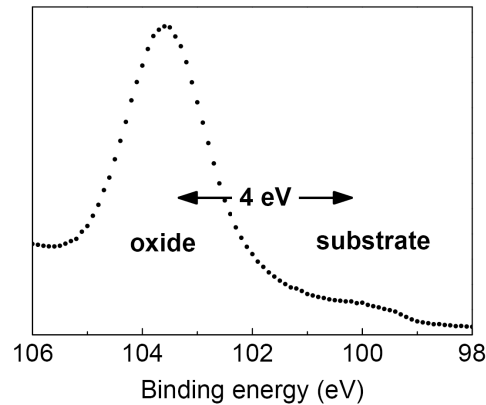
Synchrotron Results (ELETTRA)

highly p and n doped Si patterns

Threshold images of
N⁺/P⁻ and P⁺/N⁻ sample,

Si 2p
small spot pulse
counting spectra
(spectra averaged over
the field of view, 25 μ m)

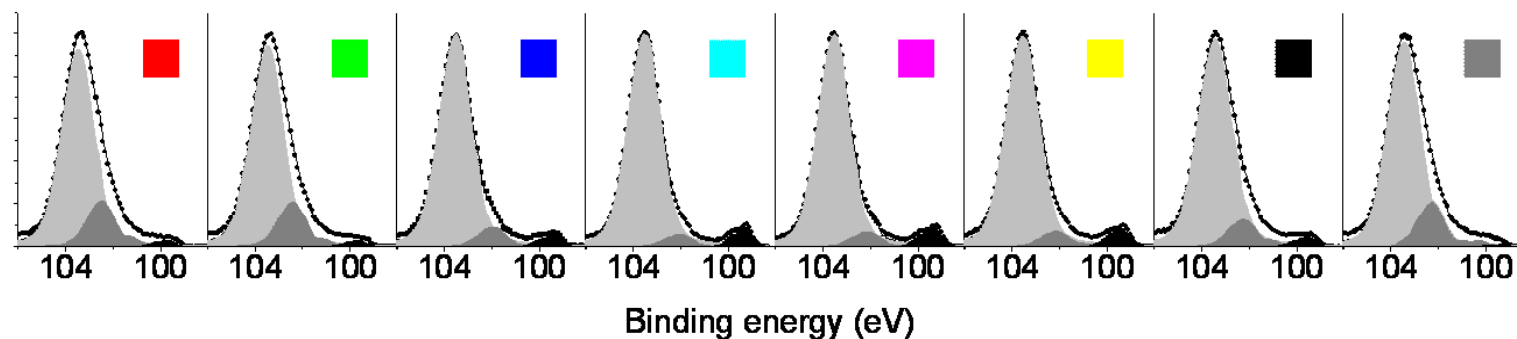
h ν = 127eV



Synchrotron Results (ELETTRA)

highly p and n doped silicon patterns $h\nu = 127\text{eV}$
Local spectra are extracted from a $400\times 400\text{nm}$
area of interest

Poster Session: We 025
M. Lavayssière



lab source results

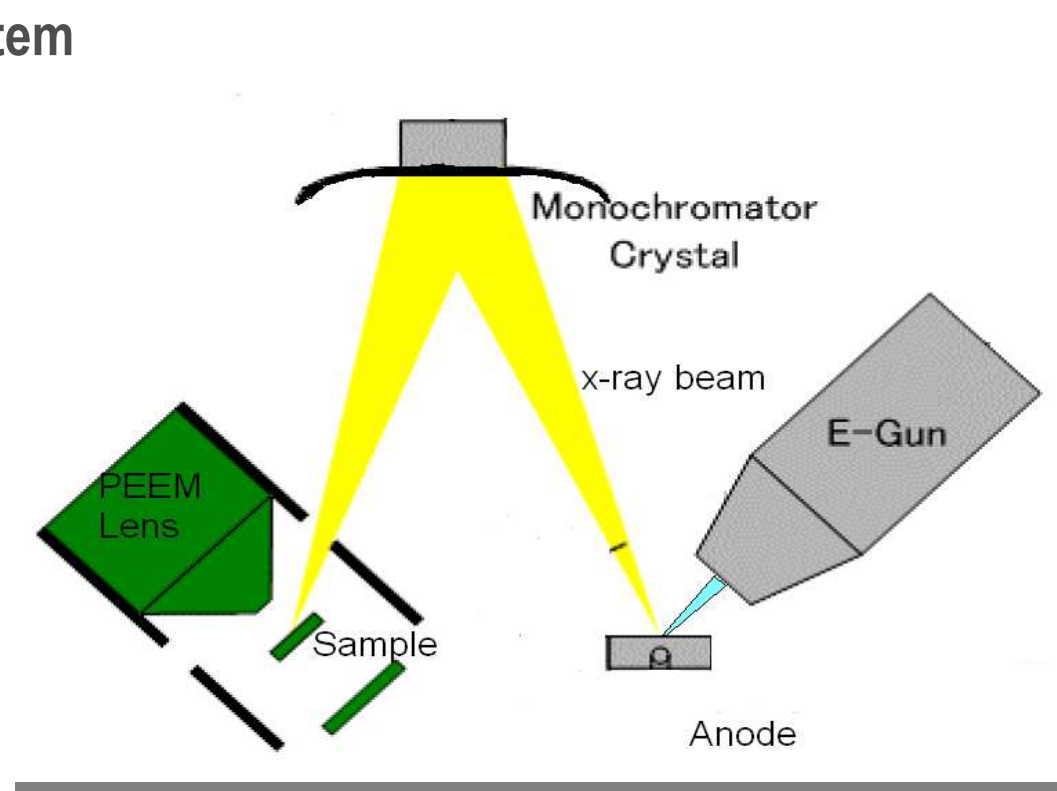
Variable x-ray spot 20-200 μm

Static x-ray spot

Imaging analyser lens system

Highest photon density

$h\nu = 1486.7\text{eV}$

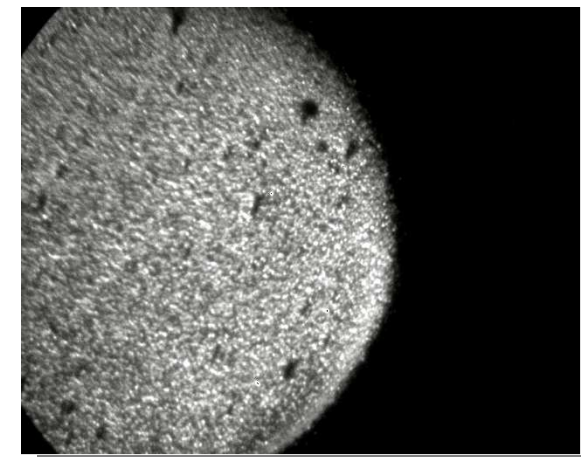
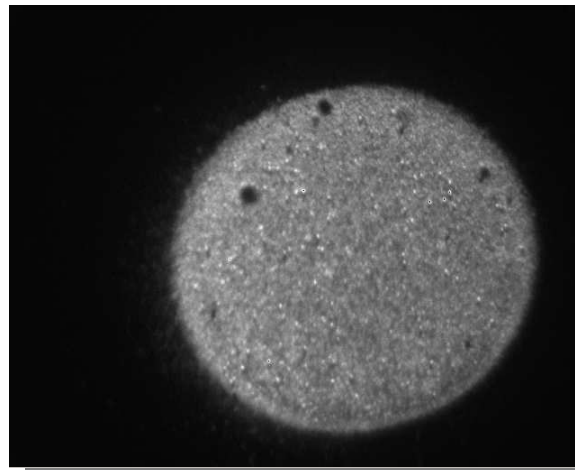
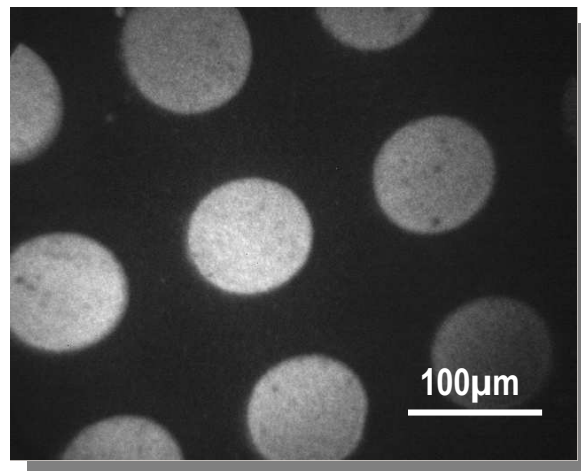


Sample: Layered Hydrogel

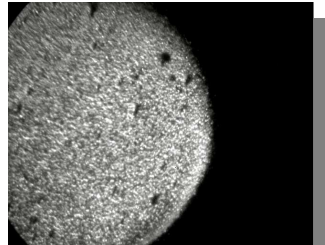
bulk: copolymer of carboxyl/hydroxyl terminated crylate/methacrylate

top (patterned) layer: amine terminated methacrylate

PEEM navigation on sample 600 μ m to 127 μ m field of view



Escher et al, submitted to *J. El. Spec. Rel. Phenom* 2009

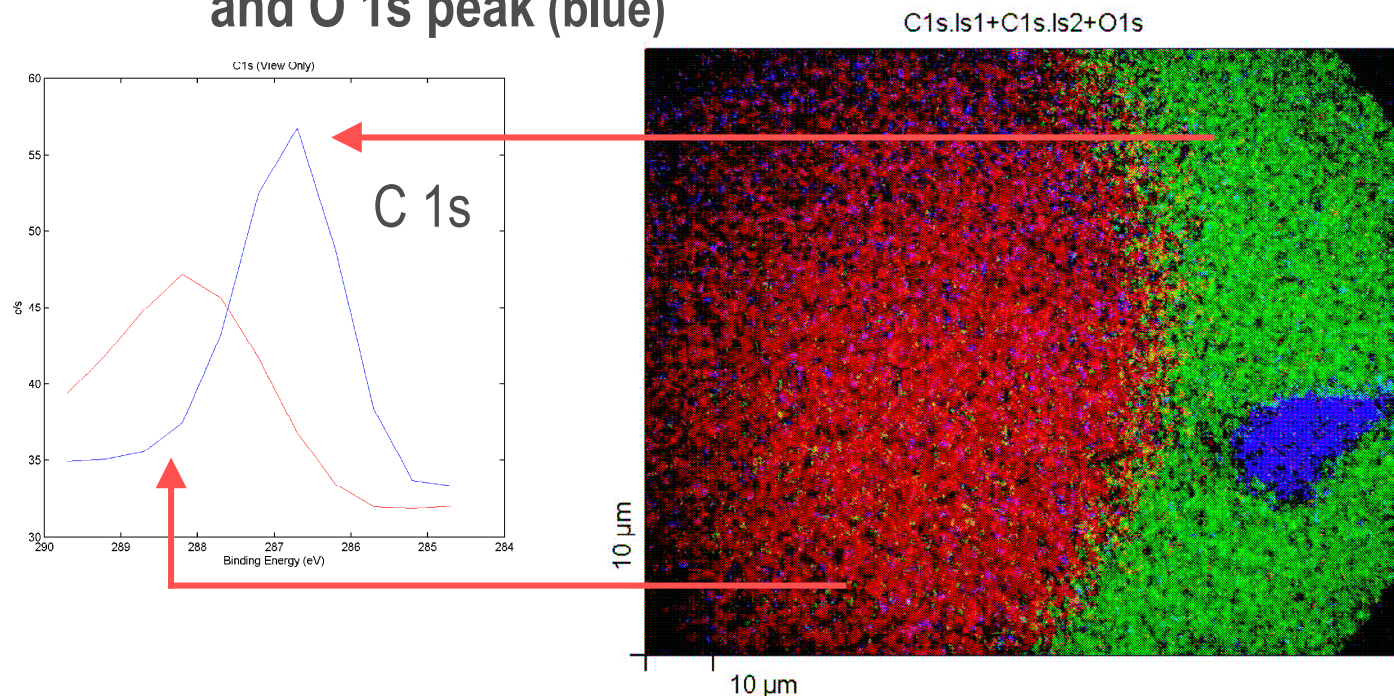


Sample: Layered Hydrogel

High resolution XPS Core level imaging

RGB overlay C 1s two modifications (red and green)

and O 1s peak (blue)



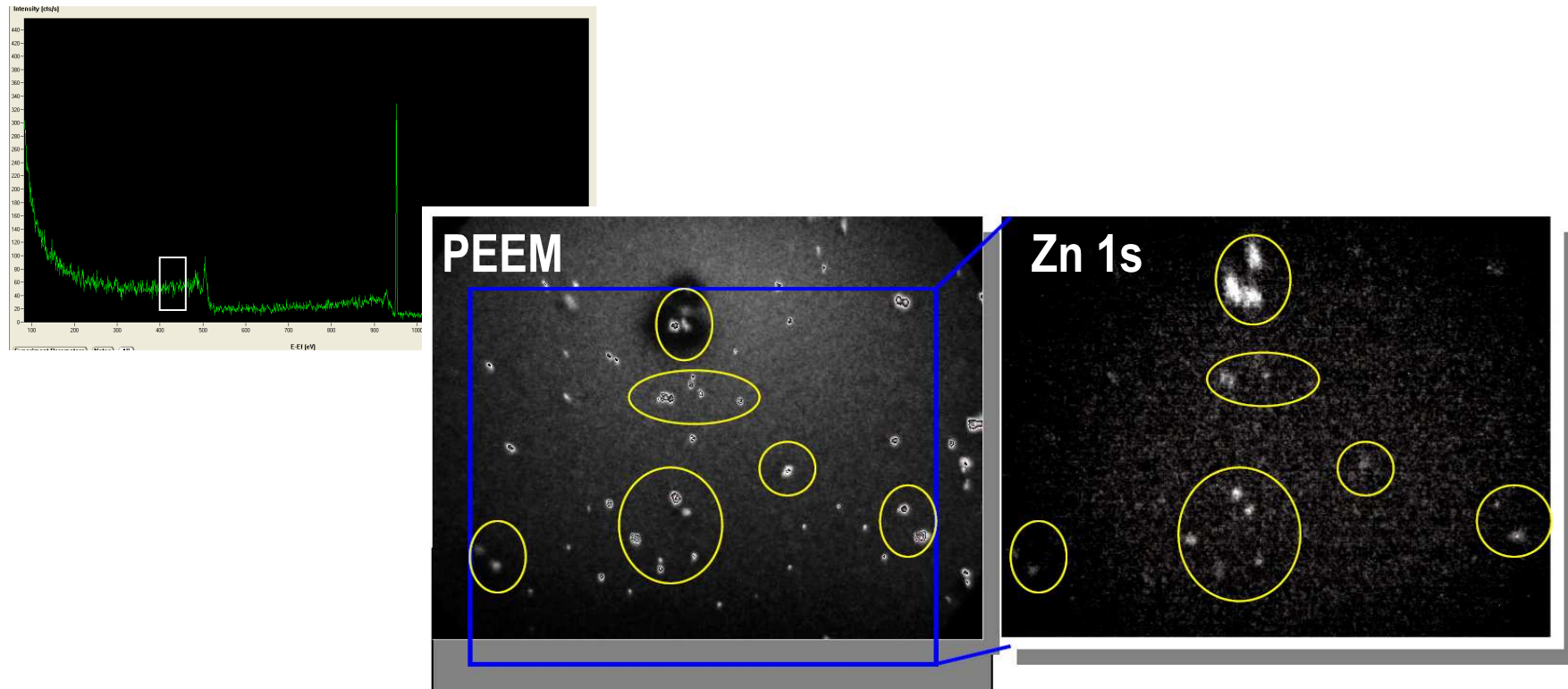
Escher et al, submitted to *J. El. Spec. Rel. Phenom* 2009

lab source results

Sample ZnO nano-particles

Small coverage of particles: weak Zn signal in small spot spectroscopy

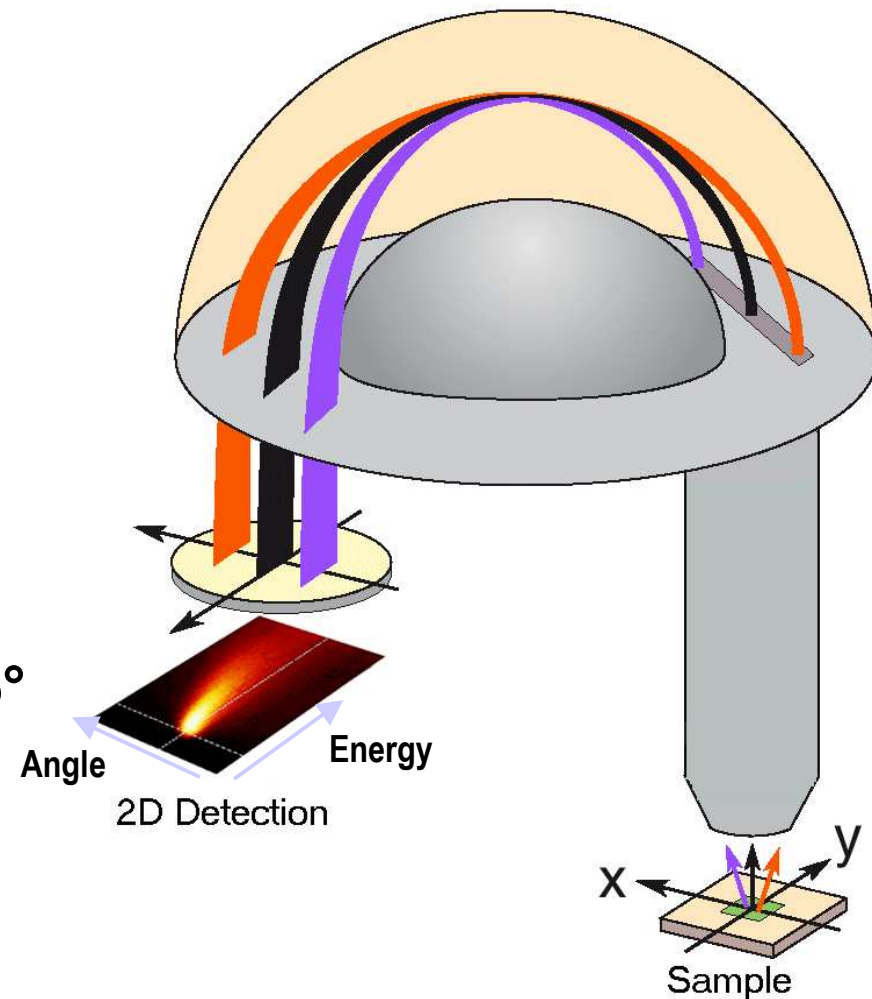
Particle size down to $<1\mu\text{m}$ are clearly visible in imaging XPS



Standard ARUPS Hemispherical Analyser with 2D detector

2 D Detectors for EA image the dispersive plane (energy spectrum) and angles simultaneously

- + Very high angular resolution
- + Very high energy resolution
- Poor control of the analysed area
- Mechanically rotated system
- Limited angular acceptance up to $\pm 15^\circ$



NanoESCA

real space imaging

Aperture at k-space image

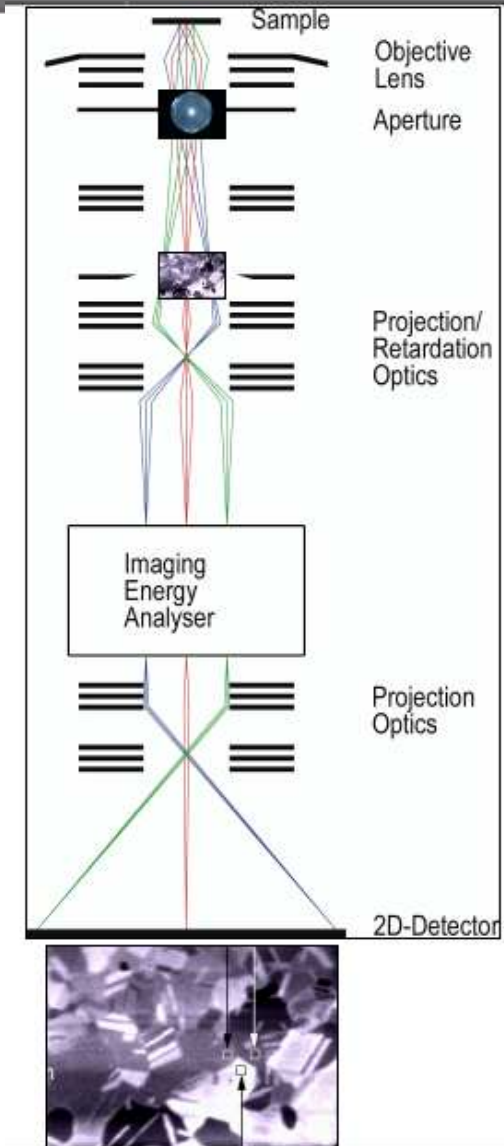
Inactive Trans Lens

Intermediate image

Image magnification

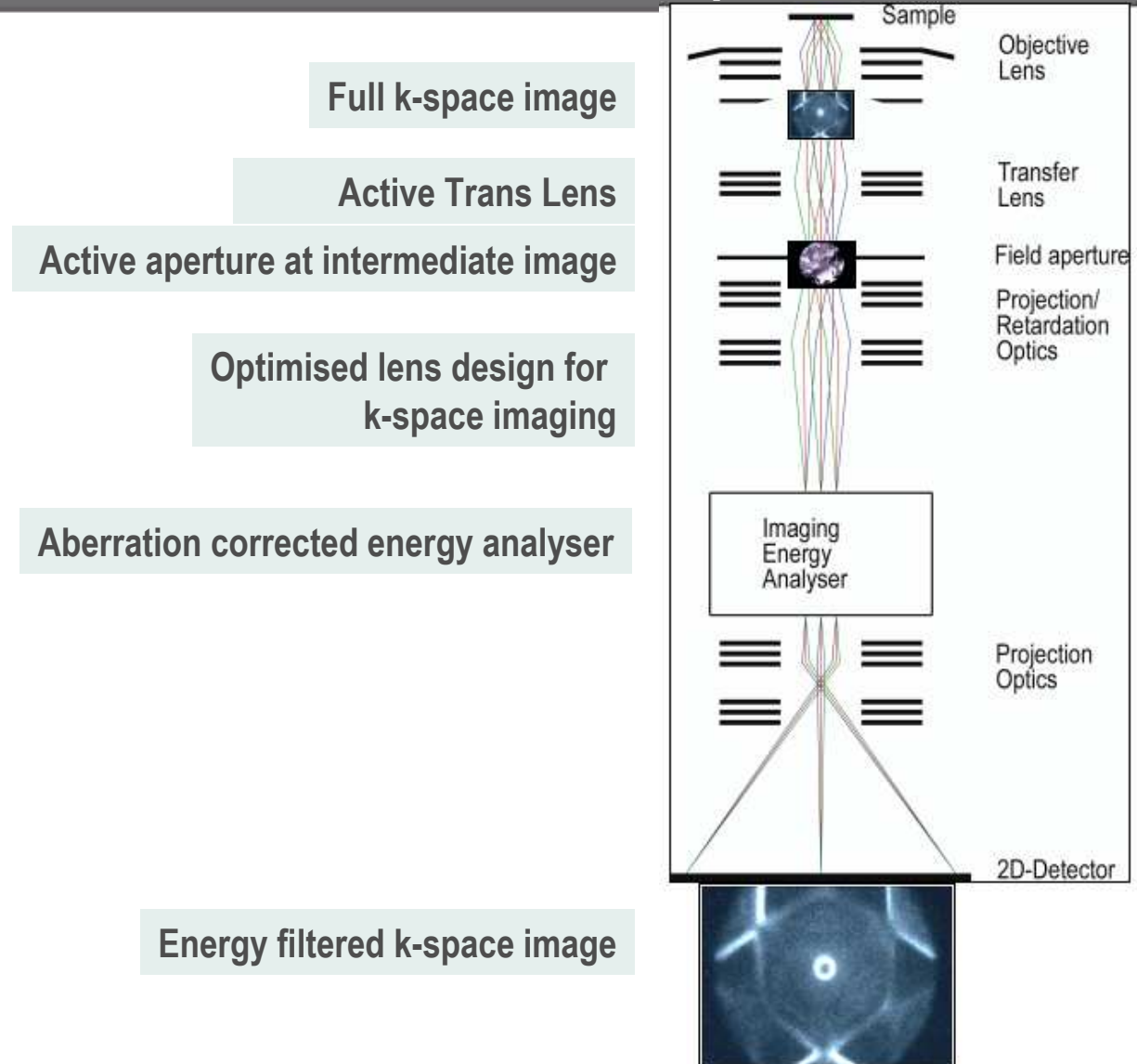
Aberration corrected energy analyser

Energy filtered real space image



NanoESCA + Transfer lense = Momentum Microscope

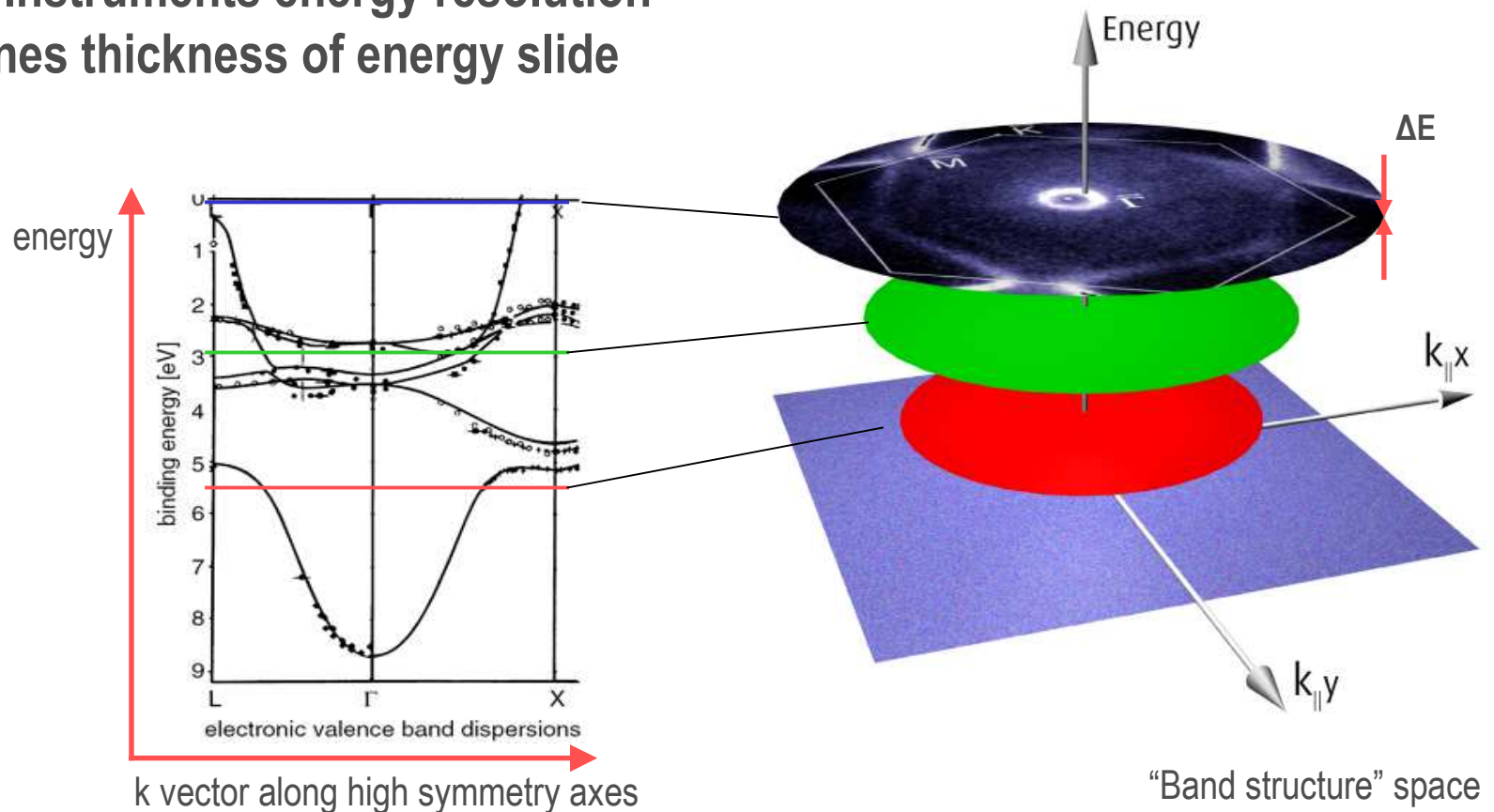
k - space imaging



NanoESCA for fast k-space mapping

Energy slices in “Band structure” space

The Instruments energy resolution defines thickness of energy slice



Single shot image of the Cu(111) Fermi Surface

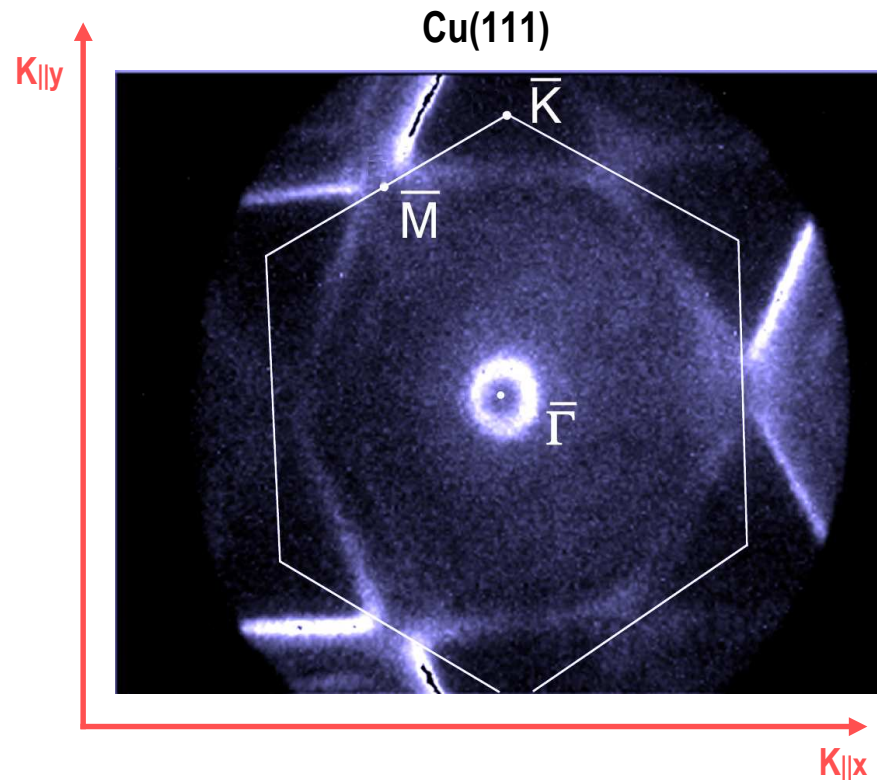
Acquisition time 60sec

He I excitation, $h\nu = 21.2 \text{ eV}$

VUV source HIS 13

With a standard ARUPS set-up it will take many hours

Large angles are very difficult to image

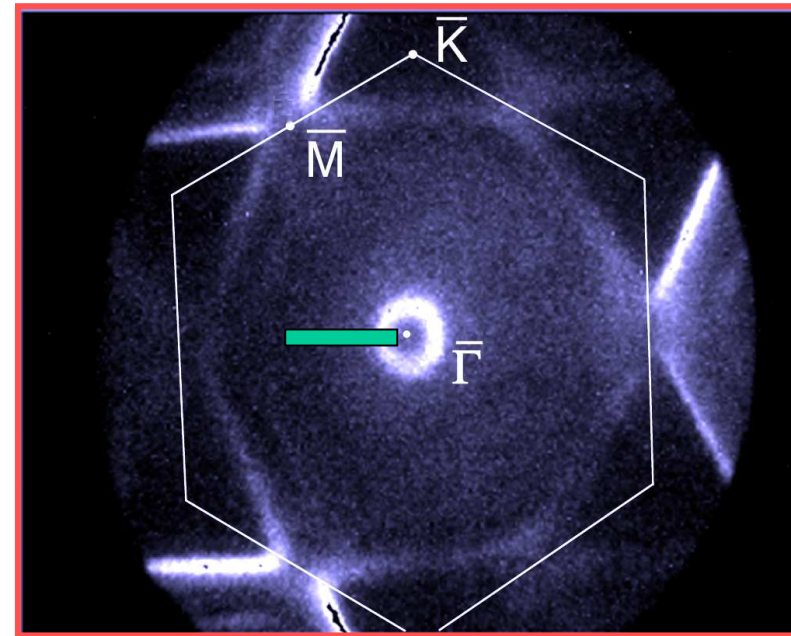
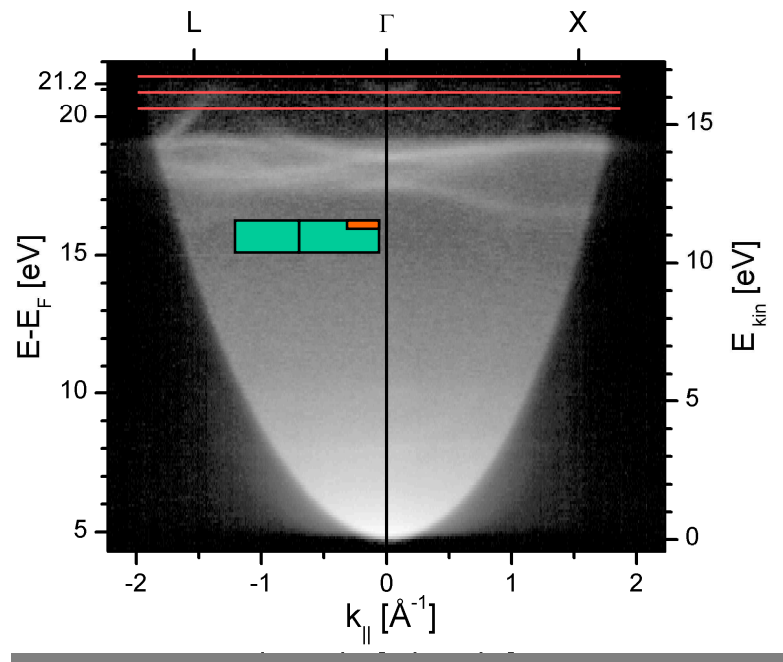


B. Krömker et al. *Rev Sci Instrum.* **79**, 2008

A scan in k-space: Cu (111)

Data $I(k_x, k_y)$ at fixed Energy

Green and red box indicate a single shot from a standard ARUPS set-up with limitations at large emission angles



B. Krömker et al. *Rev Sci Instrum.* **79**, 2008

A scan in k-space: Cu (111)

120 angular images,

ΔE per image: 50meV

60 sec per image,

2h total acq. time

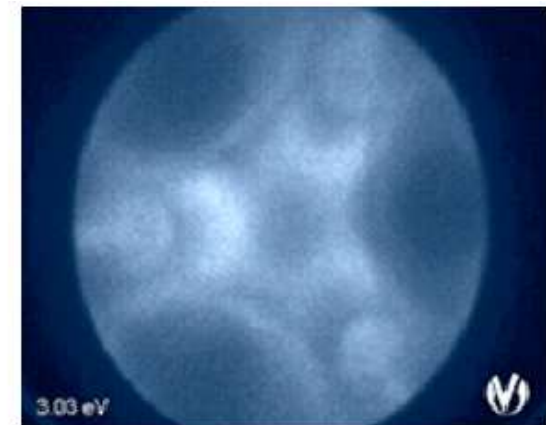
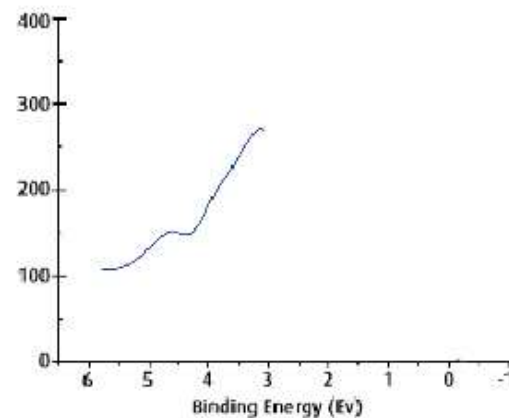
UPS Spectrum extracted

from total intensity of each

angular image

see also at

<http://www.omicron.de/nanoesca/>



120 images

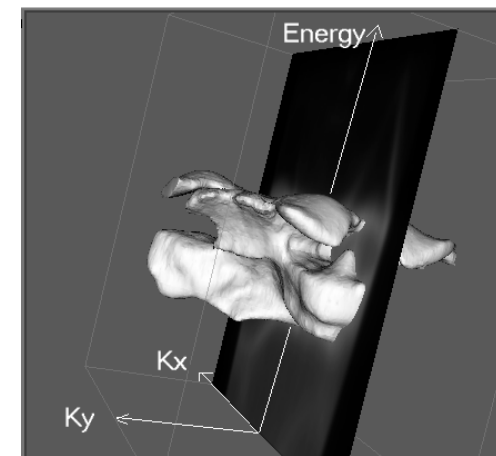
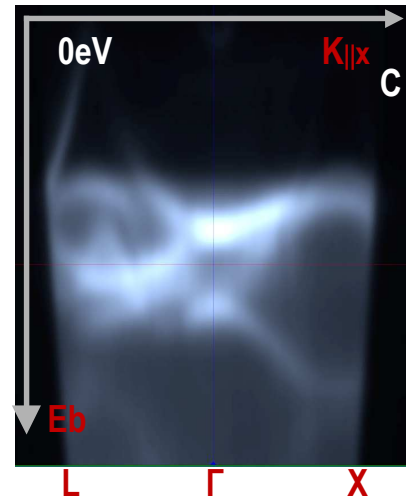
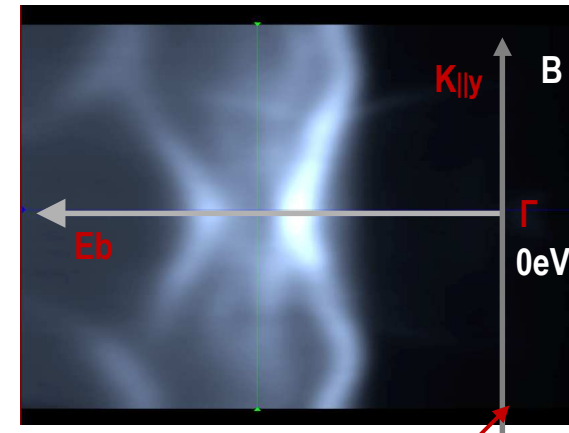
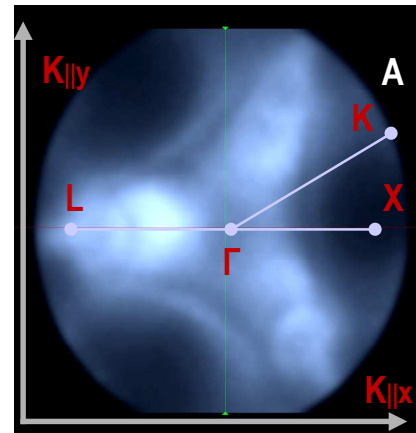
A scan in k-space: Cu (111)

Visualisation of the 120 angular images and analysis:

Cuts through image stack can be changed online.

A) Cut at constant Energy, $I(k_x, k_y)$

B) & C) Band structure cuts along along Energy axis $I(E, k_x)$ and $I(E, k_y)$



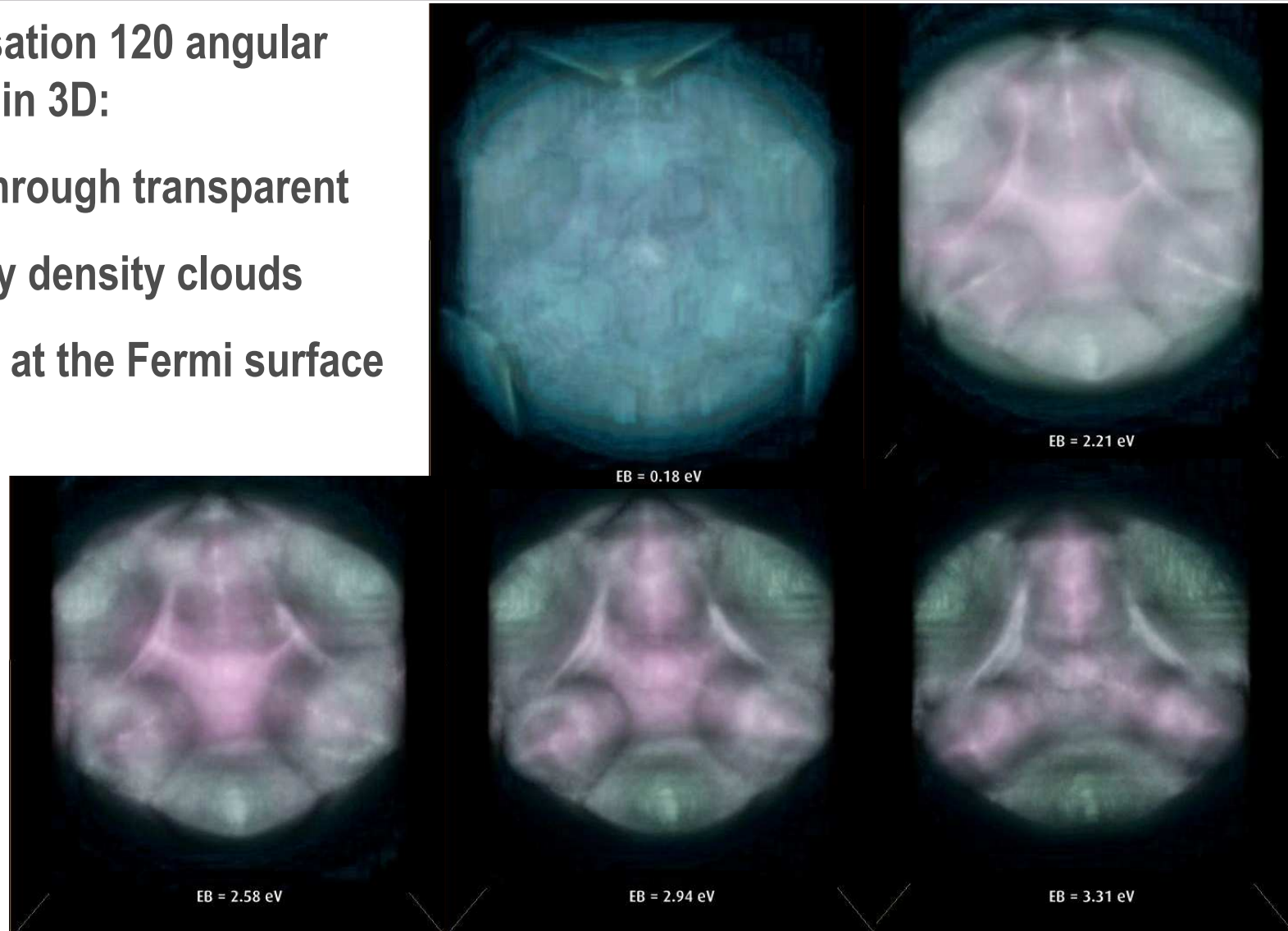
A scan in k-space: Cu (111)

Visualisation 120 angular images in 3D:

Flight through transparent

Intensity density clouds

starting at the Fermi surface



Cu (111) - Shockley Surface State

Shockley Surface State is visible with a Mercury arc lamp ($h\nu \sim 4.9$ eV)

Cu(111) surface state

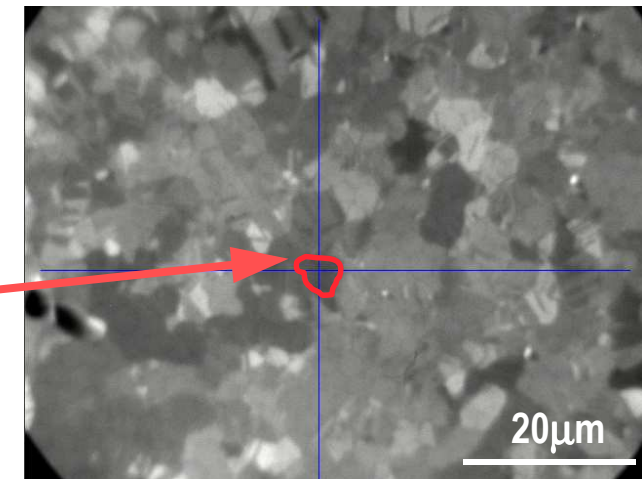
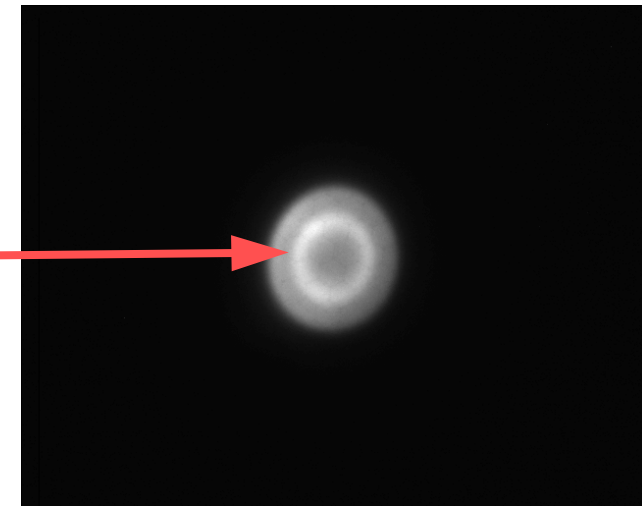
Combined real space and k-space imaging

Sample: Cu film with preferred (111) orientation and work function of 4.9 eV

O. Renault et al., Surf. Interf. Anal. 38, 375 (2006)

Single Cu grain

Cu(111)



Thank you very much for your attention!

